# **PHOTODETECTORS QUANTUM WELL**

A conventional photodetector operates by generating carriers which are produced by the absorption of a photon across the bandgap,  $E_{g}$ , of the active semiconducting region. This absorption excites an electron from the valence band to the conduction band (see Fig. 1), thereby producing a photocurrent. However, in order for this absorption to occur, the photon energy  $h\nu$  must be larger than the bandgap. This limits the useful spectral range of these detectors to the ultraviolet through near infrared region (optical wavelength  $\lambda = 0.3-5 \mu m$ ). Longer wavelengths (e.g.,  $\lambda = 8-12 \mu m$  which is an important atmospheric spectral window) require materials of very low bandgap (e.g., Hg<sub>1-x</sub>Cd<sub>x</sub>Te) that are difficult to grow, process, and fabricate into useful devices (1). Thus, it is especially difficult to make large area uniform arrays of such semiconductors, which are essential for infrared imaging applications (2). The other approach to making long wavelength arrays relies on Schottky barrier detectors which have low quantum efficiency and also require much lower operating temperatures (3).



**Figure 1.** Band structure of quantum well. Intersubband absorption between conduction band electrons levels *E*<sup>1</sup> and *E*2, or valence band hole levels  $H_1$  and  $H_2$  are shown.



**Figure 2.** Conduction band structure for a bound to contuum QWIP, showing the photoexcitation and electron transport process.

Figure 3. Geometry for QWIP photoresponse measurement.<br>For these reasons a new type of photodector (4) has been developed based on the absorption of carriers within the same band (i.e. intersubband absorption) (5). This allows the use of **QWIP PERFORMANCE** large bandgap materials which are much easier to grow and fabricate into devices, and are also far more uniform in prop- **Responsivity and Dark Current** erties than low bandgap materials. In order to use such large gap semiconductors, a *sandwich* of two different bandgap ma- The responsivity *R* of the QWIP (i.e. how much current is terials is grown on a substrate (using for example molecular generated by each incident photon) is given by (6) beam epitaxy). This creates a quantum well of the lower gap material surrounded by the larger gap barrier material, as shown in Fig. 1. Such a quantum well will have several energy levels determined by the width of the well, *L*, and the where *e* is the electronic charge,  $\eta_a$  is the absorption quantum difference in the handgaps  $\Lambda$ *E* of the two semiconductor ma. efficiency,  $p_e$  is the car difference in the bandgaps  $\Delta E_g$  of the two semiconductor ma-<br>terrier escape probability out of the well, difference in the approximate of the carrier and g is the optical gain (which is equal to the carrier terials. By controlling both of these parameters the energy and g is the optical gain (which is equal to the carrier<br>scone of the two lowest intersubbend loyels  $\Delta F - F =$  lifetime/transit time). The responsivity as a funct separation of the two lowest intersubband levels  $\Delta E = E_2$  $E_1$  can be varied over a wide range of values corresponding of length is shown for several GaAs/Al<sub>s</sub>Ga<sub>1-x</sub>As QWIPs having<br>absorption at wavelengths of  $\lambda = 3-20$   $\mu$ m. In order to create varying well widths and depth infrared optical absorption, carriers are placed in the lowest<br>infrared reached points of the responsivity,  $R$ , the detector sensitivity (i.e. the signal-<br>well will have a strong optical absorption peak at an energy<br>welr

is GaAs with  $Al_xGa_{1-x}As$  barriers since these materials are lattice matched, easy to grow, and by varying the Al composition the barrier height  $\Delta E_g$  can be readily changed. One important feature of the intersubband absorption is that due to quantum mechanical selection rules, the optical electric field must be perpendicular to the wells (5) (i.e. along the growth direction). For quick measurements, the sample is polished at an angle (e.g. 45), and the radiation is incident on this face as shown in Fig. 3, giving a substantial electric field compo nent in the normal direction. For large imaging arrays, optical gratings (7) (either periodic or random) are used to effi- **Figure 4.** Structure of a QWIP having a grating etched into the top ciently couple the light as shown in Fig. 4. Surface of the pixels.



$$
R = (e/\hbar \nu) \eta_a p_e g \tag{1}
$$

length is shown for several  $GaAs/Al<sub>x</sub>Ga<sub>1-x</sub>As QWIPs having$ 

quantum well barrier (as in the insert in Fig. 7), carriers optically excited to this continuum state will be efficiently col-**THEORY** lected and therefore the responsivity will be high. However, thermally generated carriers will also be easily collected and The most common semiconductor used for growing these wells this  $I_d$  will be large. On the other hand if  $E_2$  is bound in the





**Figure 5.** Normalized QWIP responsivity, showing the ability to easily vary the spectral response by changing the quantum well parameters. The barrier width is  $L<sub>b</sub> = 500 \text{ Å}$ , while the quantum well width  $L_w$  and the Al content  $x$  in  $Al_xGa_{1-x}As$  are given for the various samples by: A (40 Å, 0.26); B (40 Å, 0.25); C (60 Å, 0.15); D (70 Å, 0.10); E (50 Å, 0.26); and F (50 Å, 0.30/0.26). (See inserts on Figures 7–9.)

The reason that the dark current decreases the sensitivity is that it generates current noise (given for low quantum well **High Speed Response** capture probability,  $p_c$ ) as (6,10),

$$
i_n = (4eI_T g \Delta f)^{1/2} \tag{2}
$$

well (i.e. below the top of the barrier) as indicated in the lower optical gain *g* can be determined from measurements of the insert in Fig. 8, then the opposite is true: i.e. both *R* and  $I_d$  noise, and then  $p_c$  can be determined by the approximate relaare low at low voltage. The intermediate situation where the tion (6,10), (valid for small  $p_c$ )  $p_c \approx (1/g N)$ , where *N* is the excited state is exactly resonant at the top of the well or the number of wells in the QWIP. Figure 9 shows that the meaquasi-continuum case indicated in the top insert in Fig. 8 sured capture probability decreases strongly as a function of (where there are thin tunneling barriers near the top of the bias. In contrast to this the escape probability and hence the well) is near optimum, yielding a large responsivity and a low net quantum efficiency  $p = np$  incre well) is near optimum, yielding a large responsivity and a low net quantum efficiency  $\eta = \eta_a p_e$  increase strongly with bias dark current (9). as shown in Figs. 10 and 11.

Because of the very short intersubband lifetimes (1–10 ps), as well as the rapid transport of the photoexcited carriers, the where  $\Delta f$  is the bandwidth of the signal, and where  $I_T = I_d +$ <br>  $I_p$  is the total current consisting of both the dark and pho-<br>
tocurrents. In fact, by using this current noise relation the<br>
tocurrents. In fact, by using



Figure 6. Comparison of experimental (solid curves) and theoretical<br>(dashed) QWIP dark current curves at various temperatures.<br>(dashed) QWIP dark current curves at various temperatures.<br>(of Fig. 4. The insert shows the con



agram.



Figure 8. Bias-dependent (bound to bound, and bound to quasicontinuum) QWIP responsivities for samples E and F of Fig. 4. The inserts show the conduction band diagram.

tector applications.

## **Detectivity**

$$
D^* = R(A \Delta f)^{1/2} / i_n \tag{3}
$$

where *A* is the QWIP area. The detectivity is plotted as a cient for this purpose. function of temperature, *T*, in Fig. 12. for a QWIP having a<br>long wavelength cutoff (i.e. a half sensitivity wavelength) of<br> $\lambda = 10.7 \mu$ m. Note the rapid increase of *D*<sup>\*</sup> with decrease in For imaging applications (2),  $\lambda_c = 10.7 \mu$ m. Note the rapid increase of *D*\* with decrease in For imaging applications (2), the most relevant figure of merit *T* (due to the strong decrease in *i*.), with *D*\* rising from is the noise equivalent temp  $D^* = 10^{10}$  cm Hz<sup>1/2</sup>/W at 77 K to 10<sup>13</sup> cm Hz<sup>1/2</sup>/W at  $T = 35$  K.



conduction band diagram.



Figure 10. Escape probabilities vs. bias voltage for bound to continwide range of high-speed long wavelength ( $\lambda = 8-12 \mu m$ ) de-<br>norm.

From Eq. (2), we see that if  $I_d < I_p$ , then the noise and hence The detectivity  $D^*$  of the QWIP (i.e. the sensitivity of de-<br>tecting incident radiation), depends on both the responsivity<br>R and the total current noise  $i_n$  and is given by (6)<br>R and the total current noise  $i_n$  and is value of  $D^* \geq 10^{10}$  cm  $Hz^{1/2}/W$  is what is needed for excellent infrared imaging, and thus these values are more than suffi-

*T* (due to the strong decrease in *i<sub>n</sub>*), with *D*<sup>\*</sup> rising from is the noise equivalent temperature difference (3,6), *NE* $\Delta T \propto D^* = 10^{10}$  cm Hz<sup>1/2</sup>/W at 77 K to 10<sup>13</sup> cm Hz<sup>1/2</sup>/W at *T* = 35 K. 1/*D*<sup>\*</sup> which i



Figure 9. Quantum well capture probability vs. bias voltage for **Figure 11.** Quantum efficiency and escape probability vs. bias voltbound, continuum, and quasicontuum QWIPs. The inserts show the age for samples A–D. The insert shows the conduction band diagram.



0.1%), *NEAT* saturates at 60 mK for  $D^* > 10^{10}$  cm  $Hz^{1/2}/W$ . creasing the usefulness of QWIP cameras. For higher nonuniformity  $u = 10^{-4}$ ,  $NE\Delta T$  improves to 10 mK for  $D^* > 10^{10}$  cm Hz<sup>1/2</sup>/W. Thus, higher  $D^*$  is not useful when **ADVANCED TOPICS** the sensitivity is limited by nonuniformity, and in this spatial noise limit, higher uniformity leads to higher performance. We have discussed the operation of QWIPs based on conduc-<br>This is one of the main reasons that QWIPs outperform tion band electrons in quantum wells of GaAs with



detectivity. The effects of nonuniformity are included for  $u = 10^{-3}$ and 10-4

HgCdTe imaging arrays, since such low gap materials are very nonuniform due to difficulties in controlling the growth and processing. These nonuniformity problems get worse at longer wavelengths ( $\lambda > 12 \mu m$ ) for HgCdTe since the bandgap gets even smaller. In contrast, for QWIPs the semiconductors used (GaAs, AlGaAs, InP, etc.) have large gaps and are easy to grow and process into large uniform arrays. In fact large QWIP imaging arrays (128  $\times$  128) having  $\lambda = 15$  $\mu$ m have already demonstrated excellent performance (12).

Even larger GaAs/Al<sub>x</sub>Ga<sub>1-x</sub>As QWIP imaging arrays of  $128 \times 128$ ,  $256 \times 256$ , and  $640 \times 484$  pixels have been successfully demonstrated (13,14), at  $\lambda = 8-10 \mu m$  (Fig. 14) shows an image of a face with the  $640 \times 484$  array). These array sizes are much larger than is possible with HgCdTe, and QWIP imagers have achieved impressive sensitivities of  $NE\Delta T = 15$  mK. These large array sizes avoid potential thermal expansion mismatch problems between the GaAs QWIPs and the Si signal processing multiplexer to which it is bonded, **Figure 12.** Detectivity vs. temperature for a bound to continuum by thinning the QWIP array. This thinning also advanta-<br>QWIP having a cutoff wavelength of  $\lambda = 8.4 \mu$ m.<br>though nonpixel imaging has also been proposed (15 reduction in crosstalk is another advantage of QWIPs since it ence in an image. For optimum sensitivity in an imager with eliminates *blooming* (i.e. the saturation of weakly illuminated proved uniformity it is desirable for  $D^*$  and  $NFAT$  to be pixels which are near strongly illum perfect pixel uniformity it is desirable for  $D^*$  and  $NE\Delta T$  to be pixels which are near strongly illuminated pixels). Figure 14  $BLP$  However for real imaging arrays the pixel population-<br>dramatically illustrates the sen BLIP. However, for real imaging arrays the pixel nonunifor- dramatically illustrates the sensitivity and high spatial reso-<br>mity dominates the image noise (i.e. spatial noise limited) and lution of a 640  $\times$  484 QWIP ima mity dominates the image noise (i.e. spatial noise limited) and lution of a 640  $\times$  484 QWIP imager (14). Note, in particular, thus BLIP detection is not essential (3). This can be seen in the dark cool areas on the fore thus BLIP detection is *not* essential (3). This can be seen in the dark cool areas on the forehead and palm which were tou-<br>Fig. 13, where NEAT is plotted against  $D^*$  with the nonunicideal by a soda can, and the clear Fig. 13, where *NE* $\Delta T$  is plotted against  $D^*$  with the nonuni-<br>formity u as a parameter. Note that as  $D^*$  increases  $N_{T}$  and highly portable self-contained hand-held imaging camformity *u* as a parameter. Note that as  $D^*$  increases,  $N E \Delta T$  small highly portable self-contained hand-held imaging cam-<br>decreases i.e. the imaging array becomes more sensitive  $F_0r$  eras have been demonstrated with decreases, i.e. the imaging array becomes more sensitive. For eras have been demonstrated with a size and weight compa-<br>a pixel populationity of  $u = 10^{-3}$  (i.e. for a populationity of rable to home video camcorders (13), a pixel nonuniformity of  $u = 10^{-3}$  (i.e. for a nonuniformity of *rable to home video camcorders* (13), thus dramatically in-

Al*x*Ga1-*<sup>x</sup>*As barriers. In this section we will cover QWIPs based on other materials, and QWIPs using holes in the valence band.

## **Valence Band Hole QWIPs**

From Fig. 1 we can see that quantum wells in the valence band can also have intersubband absorption if doped with holes (6,16). An important difference between the conduction band electron QWIPs (*n*-QWIPs) and valence band hole QWIPs (*p*-QWIPs) is that the conduction band is nearly parabolic and thus the quantum selection rules (i.e. the requirement of a component of the optical electrical field along the crystal growth direction) holds to a good approximation. In contrast, the valence band is nonparabolic due to multiple band interactions and thus this selection rule is relaxed, allowing absorption for electric fields perpendicular to the crystal growth direction. This eliminates the need for gratings to couple the light and is thus advantageous. However, the complexity of the valence band also has a corresponding disadvantage, namely that the photoexcited carriers are scattered much more strongly and thus the carrier lifetime is **Figure 13.** Noise equivalent temperature difference as a function of shortened, lowering the gain and responsivity. This can be <sup>3</sup> seen in Fig. 15, where the responsivity of a GaAs/Al<sub>x</sub>Ga<sub>1-x</sub>As . *p*-QWIP is shown for both normal incidence as well as the



**Figure 14.** Images taken with a QWIP camera having  $640 \times 484$  pixels. Illustration courtesy of JPL (14).

of magnitude smaller than that for the usual  $n$ -QWIPs. Be-

to the maturity of this crystal system and its ease of lattice

usual 45 geometry (6). Note that both signals are compara- matched growth; however, a number of other materials have ble, due to the relaxation of the selection rules, but also that been successfully used (6). For example, the InGaAs/InP systhe magnitude of the responsivity is approximately an order tem has a similar conduction band discontinuity to  $GaAs/Al<sub>x</sub>Ga<sub>1-x</sub>As$  and QWIPs fabricated from it have similarly cause of this smaller value of *R*, the sensitivity (i.e.  $D^*$  and high performance at  $\lambda = 8 \mu m$ . By using the lattice matched *NE*<sup>I</sup>) are less than conduction band QWIPs and for this rea- quaternary lnGaAsP/InP the bandgap of the quantum well son have not been used for large-area infrared imaging can be increased and thus the response peak can be shifted arrays. to longer wavelengths. Correspondingly, by increasing the barrier height using the InP lattice matched InGaAs/InAlAs **Other Materials** system, the responsivity peak can be shifted to shorter  $\lambda$ . This  $GaAs/Al_xGa_{1-x}As$  QWIPs have received the most attention due is shown in Fig. 16 where the peak at 4.0  $\mu$ m is at the impor-<br>tant short-wave atmospheric window region. This InGaAs/



**Figure 15.** Comparison between the normal incidence and 45° re-<br> **Figure 16.** Responsivity spectrum for an InGaAs/InAlAs QWIP. sponsivity spectra for a valence band hole QWIP.



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**Figure 17.** Schematic structure for a vertically integrated two color **PHOTODIODE, AVALANCE.** See AVALANCHE DIODES. QWIP, allowing each wavelength to be individually addressed.

InAlAs/InGaAsP/InP system is thus particularly interesting since it can be used to span a very wide spectral range. In fact, by stacking the layers appropriately  $(17)$ , (see Fig. 17) a multiwavelength QWIP (with  $\lambda$  from 3-20  $\mu$ m) can be realized all on the same substrate. Additional QWIP materials systems which have been used include: GaAs/GaInP, GaAs/ AlInP, and InGaAs/GaAs.

### **SUMMARY**

QWIP imagers have demonstrated the highest pixel resolution (640  $\times$  484) in the important 8–12  $\mu$ m atmospheric window region, as well as achieving a very wide spectral range covering wavelengths from  $\lambda < 3 \mu m$  to  $\lambda > 20 \mu m$ . They have also achieved impressive sensitivities of  $NE\Delta T = 15$  mK due to their highly uniform materials and processing technologies allowed by the use of large bandgap materials. In addition, small hand-held cameras have clearly demonstrated the practicality and low cost advantages of QWIPs over low bandgap HgCdTe, and the higher sensitivity over low quantum efficiency Schottky barrier detectors. Because of these important advantages, QWIP cameras have already been successfully used (2) in medical, firefighting, military, and night surveillance applications and are expected to find many other uses in the future.

### **BIBLIOGRAPHY**

- 1. A. Sher et al., *Semicond. Sci. Technol.,* **6**: C59, 1991.
- 2. S. D. Gunapala et al., *SPIE,* **3061**: 292, 1997.
- 3. F. D. Shepard, *Infrared detectors and arrays,* SPIE, **930**: 1988.
- 4. B. F. Levine et al., *Appl. Phys. Lett.,* **50**: 1092, 1987.
- 5. L. C. West and S. J. Eglash, *Appl. Phys. Lett.,* **46**: 1156, 1985.
- 6. B. F. Levine, *J. Appl. Phys.,* **74**: R1, 1993.
- 7. J. Y. Andersson et al., *Quantum Well Intersubband Transition Physics and Devices,* H. C. Liu, B. F. Levine, and J. Y. Andersson (eds.), New York: Plenum, 1994.
- 8. M. Ershov, V. Ryzhii, and C. Hamaguchi, *Appl. Phys. Lett.,* **67**: 3147, 1995.
- 9. H. C. Liu et al., *Intersubband Transitions in Quantum Wells,* E. Rosencher, B. Vinter, and B. F. Levine (eds.), New York: Plenum, 1994.
- 10. W. A. Beck, *Appl. Phys. Lett.,* **63**: 3589, 1993.
- 11. H. C. Liu et al., *Appl. Phys. Lett.* **67**: 1594, 1995.
- 12. S. D. Gunapala et al., *IEEE Trans. Electron. Dev.,* **44**: 45, 1997.
- 13. S. D. Gunapala et al., *IEEE Trans. Electron. Dev.,* **44**: 51, 1997.
- 14. S. D. Gunapala et al., *SPIE,* **3061**: 722, 1997.
- 15. L. B. Allard et al., *Appl. Phys. Lett.,* **70**: 2784, 1997.
- 16. B. F. Levine et al., *Appl. Phys. Lett.,* **59**: 1864, 1991.
- 17. A. Köck et al., *Appl. Phys. Lett.*, **60**: 2011, 1992.

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